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Original Research Paper

Influence of Cosmic Ray Invasions and Aerosol Plasma on Powerful Atmospheric Vortices

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ABSTRACT

The Earth's atmosphere is affected by various ionizing sources. The maximum ionization of atmospheric particles by cosmic rays corresponds to the altitude of formation of tropospheric clouds. In the high-latitude troposphere for the region of the geomagnetic polar cap, in the winter period, the excitation of local cyclonic structures are observed which are accompanied with ice storms, with invasions into middle and subtropical latitudes. The time of excitation of such cyclones is about a day that is comparable with the time of excitation of tornadoes, which are generated at low latitudes. Localization of polar cyclones is not accidental. The region of the polar cap is connected with geomagnetic field lines extended into the tail of the Earth's magnetosphere. This area is open for the penetration of cosmic rays. The ionization of aerosols in the stratosphere and the upper troposphere by precipitating particles of cosmic rays enhances the vortex activity of the atmosphere. The important role of the aerosol impurity is manifested in the generation of plasma vortices and in the accumulation of energy and mass in the atmosphere by vortices during condensation of moisture. Due to the cascade character of the ionization process, the influence of cosmic radiation turns out to be non-linear and increases with increasing pollution of the atmosphere. Aperiodic electrostatic perturbations, which play a remarkable role in the genesis of vortices, are stochastically excited in plasma inhomogeneities. During the interaction of plasma vortices and Rossby vortices, a large-scale vortex structure is formed and grows.

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1. INTRODUCTION

The atmosphere of the Earth is constantly affected by various heat and ionizing sources. Particles of the atmosphere are ionized by photons, protons, ions, electrons and other particles generated by cosmic rays when nuclei are destroyed [1,2]. The average power of the solar wind injected into the magnetosphere, and, consequently, into the atmosphere, is about 10^{23} units of CGSE per day. Galactic cosmic rays are a stream of charged particles,

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32 mainly protons, in the energy range of 100 MeV – 100 GeV, penetrating the solar system
33 from interstellar space. The interaction of high-energy cosmic particles with the nuclei of the
34 components of the atmosphere leads to the formation of radionuclides. In cosmic streams,
35 protons dominate (~85%), but there are also helium nuclei (~10%), electrons (~1%) and
36 atomic nuclei with a nuclear charge up to $z \sim 30$ are observed. The process of ionization of
37 atmospheric components by energetic particles of cosmic rays is the cascade one. An
38 energetic particle of cosmic rays produces about a million or more acts of ionization in the
39 atmosphere. The maximum ionization of atmospheric particles by cosmic rays corresponds
40 to the altitude of formation of tropospheric clouds. Cloudiness and precipitation decrease
41 during the decrease in the intensity of galactic cosmic rays, and the amount of cloudiness
42 and precipitation increase after the arrival of solar cosmic rays from a flare on Earth. These
43 changes are about 10%. After the invasion of large fluxes of accelerated particles into the
44 polar regions of the Earth, a change in temperature is observed in the upper atmosphere.
45 Cosmic rays are actively involved in the formation of thunderstorm electricity.

46 Solar-terrestrial relations in the formation of atmospheric clouds and climate are nonlinear [3-
47 7]. Periods of high solar activity are usually accompanied by droughts and elevated
48 temperatures [8]. Periods of low activity in middle and northern latitudes are characterized by
49 a decrease in temperature and long snowy winters. Mechanisms of effects of the ionizing
50 flux from energetic particles of solar and galactic origin on the lower atmosphere, weather
51 and climate are presented in [9-15]. Variations in solar activity over an 11-year cycle change
52 the mean value of the solar constant (solar radiant energy flux of 1368 W/m^2) by no more
53 than 0.2%, and this change occurs in phase with the relative number of sunspots (the Wolf
54 number). Solar flares of 3 points and above are usually accompanied by the generation of
55 solar cosmic rays. There are observed synchronous changes in the density of clouds after
56 intense solar flares from solar and global meteorological data. Changes in the transparency
57 of the atmosphere under the influence of cosmic rays were detected. A noticeable increase
58 in the regeneration of cyclones off the southeastern coast of Greenland was found after the
59 onset of solar proton events [12]. The cause of changes in the temperature field may be the
60 effect of radiation effects on variations in upper clouds. The relationship between solar
61 activity and variations in the geomagnetic field is statistically traced. Fluctuations in
62 atmospheric pressure increase with periods of tens of minutes and hours during magnetic
63 storms, the intensity of the electric field of the atmosphere changes. In particular, it is shown
64 that when the Earth enters the enhanced flow of the solar wind, the pattern of distribution of
65 surface pressure changes noticeably, the instability of the troposphere increases, the
66 intensity of circulation changes, and the combination of properties of these phenomena
67 indicates the trigger mechanism of their origin. Regional studies also confirm the important
68 role of solar activity in meteorological processes [16, 17]. So, it is shown in [16] that the time
69 course of solar activity is synchronous with the distribution of dangerous meteorological
70 phenomena (in the average monthly distribution, the correlation coefficient is 0.77, in the
71 average annual distribution, the correlation coefficient is 0.82); years with a maximum of
72 dangerous meteorological phenomena are characterized by the maximum manifestations of
73 solar activity (and the drops are also synchronous). It was revealed in [17] that the wave
74 process of attenuation and increase of thunderstorm activity in the Altai Mountains is
75 subordinated to solar activity and is a link in solar-terrestrial interconnections: thunderstorm
76 minima and maxima "fall" on corresponding solar activity extremums. Ionization of the
77 atmosphere by external sources is a necessary condition for the excitation of plasma
78 vortices.

79 Polar mesoscale cyclones are an unique natural phenomenon in the atmosphere of the
80 Arctic [18, 19, 20]. They were discovered and described in the mid-twentieth century due to
81 the development of satellite sensing of the atmosphere. The applied interest in polar
82 cyclones is primarily associated with the need to predict the possible occurrence of

83 dangerous weather phenomena and their impact on the objects of economic infrastructure
84 and ships. Such structures often appear in the Arctic and have a radius not exceeding 1500
85 km. Polar winter cyclones are accompanied by ice storms. The rise time of such cyclones
86 and their lifetime is 15–25 hours, which is comparable with the time of the initiation of
87 tornadoes at low latitudes. Tornadoes are associated with hurricanes, destruction and
88 floods. Wind speed in polar winter cyclones increases up to a storm value of 15–20 m/s.
89 Invasions of polar cyclones into mid-latitudes are also observed, with abundant precipitation
90 and a sharp cooling, even to subtropics. Cyclone amplification is associated with phase
91 transitions of water. Aerosols have a marked influence on the global and regional climate
92 [21]. An important source of aerosol is volcanic eruption. Volcanic lightning is observed,
93 therefore, plasma is emitted into the atmosphere. The ionization mechanism is thermal in
94 this case, similar to friction. The kinetic energy of the movement of particles is converted into
95 ionization energy. Cosmic rays ionize particles directly in the atmosphere. Climate relations
96 are multi-parameter ones. The increasing influence of deserts on the climate is monitored.
97 This is illustrated by the example of Europe. Droughts are frequent and intensified. It affects
98 not only the Sahara desert, but also anthropogenic desert, the destruction of forests and
99 vegetation. We consider the influence of cosmic rays on the ionization of aerosol as climatic.
100 Aerosol particles play an important role in the generation of atmospheric vortices [22]. It was
101 found that aerosol particles up to 50 nanometers in diameter significantly accelerate
102 convection and precipitation [23]. At phase transitions during condensation and
103 crystallization of moisture on the aerosol particles, latent heat is released; structures of a
104 cyclonic type accumulate energy. With the mosaic cellular distribution of ionized aerosols in
105 the geomagnetic field, plasma vortices are excited on pressure gradients orthogonal to the
106 geomagnetic field [4, 6, 24, 25]. At the rotor genetic level, plasma vortices interact with
107 atmospheric vortices like particle velocity vortices.

108 **The main goal of the paper is to demonstrate the role of cosmic ray invasions and aerosol**
109 **plasma in the generation of powerful atmospheric vortices. In particular, the aim of the work**
110 **is to show that the excitation of polar cyclones in the winter period is not accidental. Such**
111 **cyclones are localized in the region of the polar cap of the geomagnetic field. Under the**
112 **pressure of the plasma of the solar wind, the magnetic force lines of the polar cap are**
113 **extended into the tail of the magnetosphere. This zone is open to the penetration of solar**
114 **and galactic cosmic rays into the geomagnetic field and precipitation into the atmosphere.**
115 **When the magnetic field lines reconnect, plasma particles in the magnetotail are accelerated**
116 **by the betatron mechanism. The precipitation of cosmic particles into the stratosphere and**
117 **upper troposphere, the ionization of aerosols, and the phase transitions of atmospheric**
118 **moisture cause nucleation of cells of a cyclonic type with an inhomogeneous mosaic**
119 **distribution of aerosol admixture in the atmosphere. Heating and ionization of particles at low**
120 **latitudes in a tornado are associated with solar photon flux and cosmic ray intrusions. An**
121 **important role in the genesis of vortices is played by aperiodic electrostatic perturbations of**
122 **an inhomogeneous plasma with an increase in the electric field inside the plasma cell.**
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124 **2. INFLUENCE OF FLUXES OF CHARGED PARTICLES AND ATMOSPHERIC** 125 **POLLUTION ON THE STRENGTHENING OF ATMOSPHERIC VORTICES IN** 126 **LOW AND HIGH LATITUDES**

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128 Polar winter cyclones are called "explosive". The wind speed increases to a storm speed,
129 about 15–20 m/s, in a time of about a few hours. Ice storm in high latitudes lasts about a
130 day. With precipitation, cyclones lose energy. Powerful polar cyclones pass into middle
131 latitudes, amplify at phase transitions of moisture, carry cooling and precipitation to
132 subtropics.

133 Changing the transparency of the atmosphere depends on the accumulation of smog of
134 natural and anthropogenic origin. Atmospheric pollution is carried by air masses at a
135 distance of thousands of kilometers. In one-dimensional geometry, the force acting on the
136 aerosol particle in the air stream is given below [26].

$$137 \quad F_s = j_n S_a p_n = N_n (v_n - v_a) S_a m_n (v_n - v_a),$$

138 where j_n - is the stream of air particles, atoms and molecules, relative to an aerosol
139 particle, S_a - is the aerosol particle frontal section, p_n - is the impulse transmitted to an
140 aerosol particle from one particle of air, m_n - is the average mass of air particles. From the
141 equation of motion of an aerosol particle $m_a (dv_a / dt) = F_s$, an estimate of the time τ of the
142 acceleration of the particle in the air flow to the speed v :

$$143 \quad -(v_n - v_a)^{-1} \Big|_0^v = \tau N_n S_a m_n / m_a.$$

144 For $v \sim 0.5 v_n$ we have $\tau = m_a / (m_n v_n N_n S_a)$. In estimating the time of capture of a particle
145 with a mass m_a and a frontal section S_a by an air stream, flow swirls when flowing around
146 the particle are not taken into account. The dynamic pressure of the flow on the particle
147 decreases with increasing particle velocity, the rise occurs with a change in acceleration.
148 Under the action of vertical pressure gradients, for example, in the vicinity of a cloudy shade,
149 aerosol impurity is pumped to cloud heights of the order of several kilometers and into the
150 stratosphere. If the lifting force and acceleration when a stream leaks onto a particle
151 exceeds the force of gravity and the acceleration of gravity, the particle rises and is carried
152 by the stream. If the particle shape is lamellar, a lifting force arises when a horizontal air flow
153 flows around the particle. The orthogonal projection of the force F_s to the frontal surface is

154 $F_l = F_s \sin \alpha$, the lifting force is $F = F_l \cos \alpha = F_s \sin(2\alpha) / 2$. At the angles of attack
155 $\alpha \sim 0, \pi / 2$, the lifting force of the horizontal flow tends to zero, but aerosol particles can rise
156 in the atmosphere at vertical pressure gradients by vertical air flows. The lifting force is
157 maximum for the angle of attack $\alpha \sim \pi / 4$. Flight properties of aerosols depend on their
158 mass, size and shape. An example of well-flying aerosols can be pollen and fine dust
159 generated by friction of transport with the road surface. For $\alpha > \pi / 2$, the particle falls. Dust
160 is blown away unevenly by the wind. The process depends on the topography of the
161 underlying surface. Moisture condensation by non-charged aerosols depends on their
162 chemical composition and surface structure. Particles can be hydrophilic and hydrophobic.
163 Ionized particles can be classified as hydrophilic, since the attraction of water molecules is
164 related to their polarization by the electric field (here, Van der Waals forces manifest
165 themselves through the phenomenon of dielectrophoresis [27]).

166 It is known that precipitations of charged particles into the atmosphere are observed mainly
167 in high latitudes of both hemispheres in auroral ovals - zones-belts surrounding the magnetic
168 poles of the Earth [28]. During the quiet Sun, on the day side, the zone boundary is $\sim 16^\circ$
169 from the magnetic pole, on the night side — $\sim 23^\circ$. During solar activity, the auroral oval
170 expands. Given the shift of the magnetic poles of the Earth from the geographic poles, the
171 boundary of the zone can reach latitude of 42° . Thus, this zone almost completely covers the
172 areas of predominantly occurring polar cyclones [18] (although complete coincidence is not
173 required, because this is not necessarily a deterministic connection, but only one of the
174 possible trigger mechanisms).

175 The formation of powerful polar winter cyclones can be associated with the eruption of
176 cosmic rays into the atmosphere in the polar cap of the geomagnetic field. Cosmic ray
177 ionization of an aerosol enhances condensation and crystallization of moisture on an
178 aerosol. The heat generated by moisture phase transitions leads to an increase in pressure
179 gradients in the cellular structure of the aerosol impurity. Plasma vortices are excited. The

180 mosaic cellular distribution of the aerosol impurity at high latitudes is formed by the
181 amplification of acoustic-gravity wave packets at wind speed gradients. During the
182 interaction of plasma vortices and Rossby vortices, an increasing cyclonic structure is
183 formed. In the genesis of vortices, aperiodic electrostatic disturbances of an inhomogeneous
184 plasma can play an important role (see below for details). In plasma flows orthogonal to the
185 geomagnetic field, an MHD electric field generator is excited, and the electric field in plasma
186 inhomogeneities is amplified.

187 The uneven mosaic distribution of smog contributes to the uneven heating of the daytime
188 atmosphere and to the excitation of atmospheric vortexes on the pressure gradients. Uneven
189 heating, ionization and condensation of moisture in aerosol cells over the ocean at low
190 latitudes provoke excitation of the tornado vortex structure. The growth of pressure gradients
191 in the cells of the uneven distribution of aerosols is also associated with condensation and
192 crystallization of moisture. This mechanism plays an important role in the excitation of polar
193 winter cyclones.

194 Aerosol impurity affects the heating of the atmosphere, is ionized by external sources,
195 condenses moisture, plays an important role in the formation of vortex structures in a
196 gyrotropic medium. Atmospheric gyrotropy is associated with the influence of the Coriolis
197 force on particles and, for charged particles in the geomagnetic field, the influence of the
198 Lorentz force.

199 As the cyclone approaching the anticyclone from the Atlantic approaches, the jet flow in the
200 area between the cyclone and the anticyclone directed northward increases. Such currents
201 carry aerosols and moisture into high latitudes. The lifetime of Aitken particles in the
202 stratosphere is years – tens of years. Their ionization by cosmic rays enhances the phase
203 transitions of atmospheric moisture on aerosols. At the same time aerosols become heavier.
204 The deposition of aerosol into the troposphere accelerates the processes of excitation of
205 plasma vortices at altitudes of cloud formation.

206 Oscillations and vortices in the atmosphere are associated with a local excess of kinetic
207 energy and free potential energy in the atmospheric layer. In atmospheric cloudiness, strong
208 electric fields and associated plasma vortices are observed, which affect the genesis of
209 atmospheric vortex structures [29-31]. Winter thunderstorms are recorded in middle and high
210 latitudes.

211 The most powerful and closest to the Earth external source of heat and ionizing radiation is
212 the Sun [32]. Particles of the atmosphere are ionized by streams of solar photons, as well as
213 photons, protons, ions, electrons and other particles, which are formed with the destruction
214 of the nuclei of atmospheric particles by cosmic rays. The process of ionization of
215 atmospheric components by energetic particles of cosmic rays is cascade, the number of
216 ionizing particles increases many times with the destruction of the atomic nucleus. For the
217 background flux of cosmic rays outside the atmosphere $j_1 \sim 0.1 \text{ cm}^{-2} \text{ sterad}^{-1} \text{ s}^{-1}$, the flux of
218 secondary particles generated by the destruction of the nuclei of atmospheric particles by
219 cosmic rays, increases by $\sim 10^6$ times – they are Auger cascade showers. The rate of
220 ionization of aerosols by cosmic ray particles is $s \sim j_2 \sigma_a N_a \sim 10^{-3} - 10^{-2} \text{ cm}^{-3} \text{ s}^{-1}$ at altitudes
221 of cloudiness for particles of Aitken, $\sigma_a \sim 10^{-10} \text{ cm}^2$, $N_a \sim 10^2 \text{ cm}^{-3}$ – are the cross-section
222 and concentration of particles. The ionization rate increases sharply with an increase in the
223 concentration of aerosol particles in the atmosphere, with an increase in ionizing fluxes
224 during solar flares and galactic fluxes. The fluxes of galactic rays are not stationary; they
225 depend on processes in stellar matter and electromagnetic fields. The ionization rate of

226 aerosol cloud changes as parameters N_a , σ_a change by several orders of magnitude. The
227 cross section of the aerosol particle is several orders of magnitude greater than the cross
228 section of atoms and molecules and increases with moisture condensation. Estimates show
229 that cosmic rays are an important source of ionization of particles in the atmosphere, even at
230 the background level of the cosmic ray flux and the concentration of Aitken particles. Cosmic
231 ray fluxes depend on the time of day, latitude and longitude, the geomagnetic field affects
232 charged particles [33,34].

233 The capture of cosmic rays by a magnetic field depends on the "rigidity" s_p , the ratio of the
234 particle momentum to the charge. "Rigidity" is associated with the Larmor radius of a particle
235 $r_B = s_p c / B$, where B is the magnetic field strength. The penetration of a particle into a
236 magnetic field depends on the field strength. It is probable that particles enter the
237 atmosphere of the Earth through the polar cap, on geomagnetic field lines stretched into the
238 tail of the magnetosphere when solar plasma fluxes flow into the geomagnetic field.
239 Atmospheric vortex activity at high latitudes during the winter period and such rare
240 phenomena as winter thunderstorms at middle and high latitudes can be associated with a
241 sharp local increase in the rate of ionization of atmospheric pollution by cosmic rays [25].

242 Intrusion of cosmic rays into the atmosphere is manifested by weak effects of ionization.
243 Ionization increases 4 times relative to ionization at sea level at an altitude of 4800 km, and it
244 increases 10 times at an altitude of 8400 km. The geomagnetic field affects the transfer of
245 plasma cosmic particles in near-Earth space. There is an association of the excitation of
246 atmospheric plasma vortices and their influence on the formation of atmospheric vortex
247 structures of cyclones and anticyclones at high latitudes (including in the polar night
248 conditions) with the excitation of energetic charged particles in the auroral zone and the
249 polar cap [35]. The transfer of humid and warm air masses to high latitudes on pressure
250 gradients with uneven heating of the atmosphere and ocean, mainly by solar photon flux,
251 and the invasion of ionizing cosmic particles into the atmosphere lead to the excitation of
252 plasma vortices [7].

253 The contribution of cosmic rays to heating is small; the energy fluxes of cosmic rays are
254 several orders of magnitude lower in comparison with the energy fluxes of solar photons.
255 The ionization of aerosol clouds by cosmic ray particles accelerates the processes of
256 moisture condensation. In phase transitions (steam, individual water molecules – water
257 droplets – ice, snow), latent heat is released, infrared radiation when the electronic shells of
258 interacting particles change during the formation of condensates. The impact of cosmic rays
259 (an ionizer of particles) on phase transitions in the atmosphere with a non-uniform spatial
260 distribution of aerosols (aerosol clouds) during the condensation of moisture by aerosols
261 enhances the vortex activity of ionized aerosol impurities [25].

262 The generation and stability of plasma vortices is influenced by electromagnetic radiation
263 and other factors, for example, acoustic-gravity waves [36-39], the electric and magnetic
264 fields of various sources, including man-made, associated with human activity. With the
265 accumulation of mass by a moving plasma vortex, a state of weakly stable holding by the
266 vortex of the cloud mass in the field of gravity can be achieved. In such a state, even weak
267 external influences on the structure of the plasma vortex can stimulate precipitation and the
268 extinction of the vortex, or vice versa, the strengthening of the vortex [40].

269 When ionized aerosols move in fields of pressure gradients orthogonal to the geomagnetic
270 field, electric fields and plasma vortex structures are excited in cellular fields of aerosols.
271 There is a transfer of thermal energy of non-uniform heating into a plasma vortex motion in

272 crossed fields, an electric field of a vortex and a geomagnetic field. Ions of different mass
273 may be involved in this movement. The equation does not depend on the mass of the ions.

274 Theoretical studies of the spatio-temporal evolution of coherent structures (such as
275 cyclones, anticyclones, Rossby vortices, etc.) received a new impetus after the discovery of
276 the analogy between the Hasegawa-Mima equation [41,42] and the barotropic vorticity
277 equation, which is used to describe large-scale vortex flows in the atmosphere and ocean.

278 In our case, deriving the conservation equation for the plasma vortex in the presence of a
279 frontal temperature jump ($\partial \ln T / \partial y \gg \partial \ln T / \partial x$), we can obtain an equation of the type of
280 the generalized Hasegawa – Mima equation [43]:

$$281 \quad \frac{\partial}{\partial x} [-v_{dw} \ln(\Delta \Phi) + \frac{e}{T} v_{dw} \Phi + \frac{e}{M \Omega_{0i}} \frac{e}{T} \frac{\partial \ln T}{\partial y} \frac{\Phi^2}{2}] + \frac{e}{M \Omega_{0i}} J(\Phi, \ln(\Delta \Phi)) = 0, \quad (1)$$

282 where $J(\Phi, \ln(\Delta \Phi)) = \Phi_x (\ln(\Delta \Phi))_y - \Phi_y (\ln(\Delta \Phi))_x$, v_{dw} - is the vortex drift velocity, M - is the
283 ion mass, Ω_{0i} - is the ion cyclotron frequency, the axis Z is directed along the external
284 magnetic field, Φ - is the electric field potential. The components of the drift velocity along
285 the X and Y axes are determined by the formulas, respectively:
286 $u = -e\Phi_y / (M \Omega_{0i})$, $v = e\Phi_x / (M \Omega_{0i})$, where $\Omega' = v_x - u_y = e(\Phi_{xx} + \Phi_{yy}) / (M \Omega_{0i})$ - is the drift
287 rotor. From equation (1) it follows

$$288 \quad \frac{\partial}{\partial x} [-v_D \ln(\Delta \Phi) + \frac{e}{T} v_D \Phi + \frac{ce}{BT} \frac{\partial \ln T}{\partial y} \frac{\Phi^2}{2}] + \frac{c}{B} J(\Phi, \ln(\Delta \Phi)) = 0. \quad (2)$$

289 Changes in the geomagnetic field can lead to disturbances and the breakdown of the plasma
290 vortex. The electric field of the plasma vortex:

$$291 \quad \mathbf{E} = \ln(N / N_0) (\nabla T) / e + (T / e) \nabla \ln N,$$

292 the rotational velocity of plasma particles in crossed fields:

$$293 \quad \mathbf{v}_d = [c / (eB^2)] \{ \ln(N / N_0) [\nabla T \times \mathbf{B}] + T [\nabla \ln N \times \mathbf{B}] \},$$

294 the particle concentration, the temperature of the electronic component T , and the potential
295 of the electric field are related by the Boltzmann distribution [43]. The energy density of the
296 electric field of the vortex W_1 is $W_1 = (8\pi e^2)^{-1} \{ \ln(N / N_0) \nabla T + T \nabla \ln N \}^2$, where the
297 concentration distribution of particles in the plasma vortex structure is given in the form
298 $N = N_0 \exp(e\Phi / T(y))$, T - is the electronic component temperature, N_0 - is the particle
299 concentration for $\Phi = 0$.

300 Plasma instabilities of inhomogeneous plasma provoke the development of nonlinear
301 processes. Different types of instabilities can be observed in inhomogeneous plasma of
302 ionized smog [44]. So, electrostatic oscillations can be excited, for example, on drift gradient
303 plasma instabilities. If the density gradient of particles is directed against gravity $\nabla N \uparrow \downarrow \mathbf{g}$,
304 the cloud of smog is unstable, where N is the concentration of particles, \mathbf{g} is the acceleration
305 of the gravitational field. For this type of instability, the main role is played by heavy particles
306 (ions), the contribution of the light electronic component is small. In ionized cloudiness,
307 impurity instability can be excited if the density gradient of an impurity of heavy cold ions is
308 directed against the density gradient of "hot" background plasma. In layers with a plasma
309 density gradient directed against gravity, it is possible the development of hydrodynamic
310 instability with the growth increment of the order $\gamma \sim \sqrt{g \partial \ln(N) / \partial y}$, where $\partial \ln(N) / \partial y$ is the
311 vertical relative gradient of plasma density. Drift instabilities affect polar cyclones as they
312 propagate to low latitudes. In non-uniform magnetized plasma, gravitational-dissipative
313 instability can develop. The development of instability is most likely in low and middle

314 latitudes. Frequency and growth increment of free oscillations are respectively detailed in the
 315 equation given below [44]:

$$316 \quad \omega = \omega_{ni} + g \kappa / \omega_{ni}, \quad \gamma = (\gamma_0 / \omega_{ni})^2 z_i v, \quad$$

317 where N – is the concentration of plasma, $\omega_{ni} = k_y \kappa c T_i / e_i B$, $\gamma_0 = \sqrt{g \kappa}$, $\kappa = \partial \ln N / \partial x$, c – is
 318 the speed of light, $z_i = k_{\perp}^2 T_i / (m_i \Omega_i^2)$; $(T_i, e_i, m_i, \Omega_i)$ – are accordingly the temperature, charge,
 319 mass and ion cyclotron frequency, (k_x, k_y) – are the components of the wave vector of
 320 perturbations perpendicular to the geomagnetic field and along the axis \mathbf{y} , $\mathbf{x} \uparrow \downarrow \mathbf{g}$. An
 321 orthogonal coordinate system with a \mathbf{z} axis parallel to an external magnetic field is selected,
 322 $\mathbf{z} \parallel \mathbf{B}$. For the region of the geomagnetic equator, it is assumed that $\mathbf{g} \perp \mathbf{B}$. Below the
 323 maximum of the ionospheric layer F , where $\nabla N \uparrow \downarrow \mathbf{g}$, the development of the gradient
 324 instability of a magnetized ionospheric plasma in a gravitational field leads to the
 325 appearance of plasma inhomogeneities and scattering of radio waves. Gravitational-
 326 dissipative instability at low and middle latitudes may contribute to the separation of aerosol
 327 plasma in atmospheric cloudiness and the excitation of plasma vortices in the
 328 inhomogeneous plasma structure.

329 Using [45], we obtain an analytical solution for the dielectric constant of a nonstationary
 330 inhomogeneous plasma. From the Vlasov equation for the electron distribution function

$$331 \quad \left\{ \frac{\partial}{\partial t} + \mathbf{v} \nabla + \left(\frac{e \mathbf{E}}{m} + [\mathbf{v} \Omega] \right) \nabla_{\mathbf{v}} \right\} f = 0$$

332 for the case of one-dimensional distribution of particles in velocity, without magnetic field
 333 $\Omega = 0$, in linear approximation $f \rightarrow f_0 + f$ for electric field disturbances, we have:

$$334 \quad f = - \int_0^{\infty} \exp(-v \tau \nabla) \left(\frac{e \mathbf{E}}{m} \nabla_{\mathbf{v}} f_0(t - \tau) \right) d \tau,$$

335 where the retarded potential $\exp(-v \tau \nabla)$ for cold plasma $f_0(v) = \delta(v)$ leads only to
 336 convergent integrals. We assume that electrostatic perturbations are potential $\mathbf{E} = \nabla \phi$ and
 337 proceed to the spectrum in the one-dimensional case $\phi \sim \int \phi_{k, \omega} \exp(ikz - i\omega t) dk d\omega$. Then the
 338 dispersion equation of a cold inhomogeneous plasma follows from the Poisson equation
 339 $\text{div} \mathbf{E} = 4\pi e \int f d\mathbf{v}$. Let us assign the distribution function of particles nonstationary and non-
 340 uniform along the \mathbf{z} axis in the form $f_0 = N_0 \exp(ct) \exp(-z^2 / b^2) \delta(v)$ and for $c < 0$ we obtain
 341 the dispersion equation of electrostatic disturbances of the electronic component of the
 342 plasma inhomogeneity in operator form

$$343 \quad - \int_{-\infty}^{\infty} k^2 e^{ikz - i\omega t} \phi(k, \omega) dk d\omega = - \frac{4\pi e^2}{m} N_0 e^{ct} \int_{-\infty}^{\infty} d\mathbf{v} \int_0^{\infty} d\tau e^{-c\tau} \exp(-v \tau \nabla) \int_{-\infty}^{\infty} ik \exp[ikz - i\omega(t - \tau)] \times$$

$$344 \quad \phi(k, \omega) dk d\omega \int_{-\infty}^{\infty} dk' e^{ik'z} N_0(k') \frac{\partial \delta(v)}{\partial v}. \quad (3)$$

345 For small v we have $\exp(-v \tau \nabla) \sim 1 - v \tau \nabla$ and, using the delta function representation as

$$346 \quad \delta(x, \alpha) = \frac{\alpha}{\pi^{1/2}} \exp(-x^2 \alpha^2), \quad \lim_{\alpha \rightarrow \infty} \delta(x, \alpha) = 0, \quad x \neq 0,$$

347 taking into account the orthogonality of the
 348 spectral components, we obtain the dispersion equation of electrostatic perturbations from
 expression (3):

$$349 \quad \varepsilon = 1 + \omega_p^2 \exp \left(ct - \frac{z^2}{b^2} \right) \left[1 + \frac{2zi}{kb^2} \right] \frac{(c^2 - \omega^2 + 2i\omega c)}{(c^2 + \omega^2)^2} = 0. \quad (4)$$

350 Aperiodically increasing electrostatic disturbances manifest themselves in a simple
 351 approximation for a cold plasma that is non-uniform along the z -axis with a particle
 352 distribution function

$$353 \quad f_0(z, v_z) = N_0 \exp(-z^2 / b^2) \delta(v_z),$$

$$354 \quad k = \frac{(2zi\omega_p^2 / b^2 \omega^2) \exp(-z^2 / b^2)}{1 - (\omega_p^2 / \omega^2) \exp(-z^2 / b^2)}, \quad (5)$$

355 where k – is magnitude of the spatial vector of electrostatic disturbances, $\omega_p^2 = 4\pi N_0 e^2 / m$ –
 356 is the square of the plasma frequency of electrons for the center of inhomogeneity. Provided
 357 $1 - (\omega_p^2 / \omega^2) \exp(-z^2 / b^2) < 0$ at frequencies less than the local plasma frequency, the electric
 358 fields increase.

359 Aperiodic electric fields are excited in plasma inhomogeneities. An electric field component
 360 appears, parallel to the power lines of the geomagnetic field, which is important for the
 361 formation of lightning channels. The formation of the cellular structure of inhomogeneous
 362 plasma is stochastically deterministic. Plasma vortices are excited in plasma with a magnetic
 363 field in the cells. The formation of the cellular structure is associated with the electrostatic
 364 instability of inhomogeneous plasma, an aperiodic increase of the electric field in the plasma
 365 inhomogeneity. In extremums at $\partial k / \partial z = 0$, the growth of aperiodic electric fields is not
 366 monotonous. Aperiodic electric fields along geomagnetic field lines enhance the interaction
 367 of plasma vortices in a geomagnetic field tube. This leads to an acceleration of the growth of
 368 the atmospheric vortex structure during the interaction of plasma vortices with Rossby
 369 vortices at the rotor level as the particle velocity vortices. The geomagnetic field affects the
 370 localization, excitation and interaction of plasma vortices. In the polar zone, where the
 371 inclination of the geomagnetic field lines relative to the horizontal approaches $\pi / 2$, a local
 372 explosive cyclone is excited. Part of the energy of atmospheric vortex structures is
 373 generated by plasma vortices, while the excitation and interaction of plasma vortices
 374 depends on the mechanisms of generation of electric fields. A cellular MHD generator and
 375 associated plasma vortices are excited on pressure gradients orthogonal to the geomagnetic
 376 field. In inhomogeneous plasma along the geomagnetic field, aperiodic electric fields are
 377 excited, accelerating the interaction of plasma vortices. Atmospheric vortex structure in
 378 aerosol plasma increases. Electromagnetic plasma processes play an important role in the
 379 mechanisms of localization, excitation and amplification of atmospheric vortex structures -
 380 cyclones, anticyclones, and tornadoes in a double gyrotropy medium. In a polar atmosphere,
 381 the cyclone structure acquires an explosive character like a tornado at low latitudes. This is
 382 facilitated by the generation of electric fields in geomagnetic power tubes with aerosol
 383 plasma.

384 The formation of mosaic cellular distributions of aerosol impurities at high latitudes is
 385 associated with the propagation and amplification of acoustic-gravity wave packets at wind
 386 speed gradients. We consider the effect of wind shifts on wave propagation. Let the
 387 horizontal directional velocity depend only on the vertical coordinate z . In the coordinate
 388 system moving with the constant wind speed $U_0(z)$ along the x axis, the system of linearized
 389 hydrodynamic equations of motion for a material point can be written as

$$390 \quad \frac{\partial u}{\partial t} + w \frac{\partial U_0}{\partial z} = - \frac{1}{\rho_0} \frac{\partial p}{\partial x}, \quad \frac{\partial w}{\partial t} = - \frac{1}{\rho_0} \frac{\partial p}{\partial z} - g \frac{\rho}{\rho_0},$$

391 where u, w - are small perturbations of horizontal and vertical speed, p, ρ - are pressure
 392 and density disturbances, ρ_0 - is density, g - is acceleration of gravity. The transition to a
 393 fixed laboratory coordinate system from a moving system is associated with the Doppler

394 frequency shift $\pm k U_0$ (the sign depends on the directions U_0 and k). In the linear
 395 approximation, the continuity equation for an incompressible medium when moving in a
 396 plane x, z is represented as $\frac{\partial \rho}{\partial t} + w \frac{\partial \rho_0}{\partial z} = 0$. Differentiating with respect to t the equation for

397 w and the equation of incompressibility $\frac{\partial u}{\partial x} + \frac{\partial w}{\partial z} = 0$ and substituting

398 $(w, p) \sim (w(z), p(z)) \exp(ikx - i\sigma t)$, $w = W \exp(-c_1 z)$, $c_1 = \left(\frac{\partial \ln \rho_0}{\partial z} + \frac{k}{\sigma} \frac{\partial U_0}{\partial z} \right) / 2$ for the case

399 $\frac{\partial U_0}{\partial z} = const$, we obtain

400
$$\frac{\partial^2 W}{\partial z^2} + \left\{ -\frac{1}{4} \left(\frac{\partial \ln \rho_0}{\partial z} + \frac{k}{\sigma} \frac{\partial U_0}{\partial z} \right)^2 + \right.$$

401
$$\left. + \frac{k^2}{\sigma^2} (N^2 - \sigma^2) + \frac{k}{\sigma} \frac{\partial \ln \rho_0}{\partial z} \frac{\partial U_0}{\partial z} \right\} W = 0, \quad (6)$$

402 where the sign of the expression in braces $\{ \dots \} = s_1$ determines the type of equation. If
 403 $s_1 > 0$, equation (6) is the equation of oscillations, and when $s_1 < 0$ - the equation of damped
 404 or increasing perturbations:

405
$$\partial^2 W / \partial z^2 - a^2 W = 0, \quad W = c_2 \exp(\pm a z);$$

406
$$\partial^2 W / \partial z^2 + a^2 W = 0, \quad W = c_2 \exp(\pm i a z),$$

407 where the arbitrary constant c_2 can be determined, for example, from boundary conditions at
 408 $z = 0$. The transition to the laboratory coordinate system is made by replacing $\sigma \rightarrow \sigma \pm k U_0$.

409 From these equations it can be seen that the wind shear significantly affects the wave
 410 disturbance. The relationship of vertical and horizontal movements is determined by the
 411 complex ratio of environmental parameters. In the vertical plane, the wave number for the
 412 equation of oscillations turns out to depend on the parameters of the inhomogeneous
 413 medium, as does the amplitude of oscillations. For $s_1 < 0$ the buoyancy wave fade out,
 414 leaving the channel where the horizontal wind speed is constant. Depending on the ratio
 415 c_1, s_1 , an increase in disturbances at the wind discontinuity is possible. With a decrease in
 416 the density of the atmosphere $\partial \rho_0 / \partial z < 0$, the amplitude of the waves can also increase. On
 417 the gradients of wind speed, acoustic-gravity wave packets are amplified. The effect of
 418 mosaic cellular distributions of atmospheric pollution on the enhancement of vortex
 419 structures, jet flows and turbulence is theoretically tracked. In the approximation of a
 420 compressible medium, the square of the Brunt-Väisäl frequency is [46]:

421
$$N^2 = -g \left[\frac{\partial \ln \rho_0}{\partial z} + \frac{g}{c_s^2} \right], \text{ where } c_s - \text{ is the speed of sound.}$$

422 The pressure gradients in the cyclonic vortex cell are directed from the center of the cell to
 423 the periphery. This is connected with the condensation of moisture, the growth of the optical
 424 thickness of the cell (the cloud) and with increasing pressure gradients in the vicinity of the
 425 cloudy shade. When a moist air mass moves from the periphery of the aerosol cell to its
 426 center, water molecules pass into condensate, and latent heat is released. The
 427 concentration of water molecules in the formation of condensate gradually decreases, and
 428 therefore the rate of formation of condensate decreases from the periphery of the aerosol
 429 cell to the center. The effect of aerosol plasma on cell formation of a cyclonic type is clearly
 430 manifested in the structure of a tornado. In anticyclonic cells, pressure gradients are directed
 431 toward the center. As a consequence, the above considered movement of humid air and the

432 formation of condensate in cells of anticyclonic type does not occur, in contrast to cells of
433 cyclonic type. In the large-scale vortex structure of the anticyclone, local formation of cells of
434 a cyclonic type is possible.

435 The formation of polar winter cyclones is associated with a mosaic cellular distribution of
436 aerosols. Such spatial distributions of the aerosol impurity can occur when acoustic-gravity
437 waves are amplified at the wind velocity gradients in the jet streams. Jet streams directed to
438 high latitudes are formed at the boundary of the anticyclone and the cyclone approaching it.
439 Anticyclone accumulates pollution. The influence of anthropogenic pollution on the
440 cyclogenesis of the atmosphere is increasing. The influence of cosmic rays on atmospheric
441 processes in the polar cap is intensified, since this region of the geomagnetic field is open
442 for cosmic rays.

443 Global jet streams circulate around the Earth: two polar, two subtropical and one equatorial
444 [47,48]. The width of the flows horizontally is hundreds of kilometers, vertically - less than 5
445 km, height - about 11 km. Under the influence of atmospheric vortex structures, jet streams
446 become sinuous. The jet flow directed northward forms in the region of the approach of the
447 Atlantic cyclone and anticyclone. The jet flow transfers pollution and moist cyclonic air into
448 high latitudes, and under its pressure the polar flow acquires a bend. A polar cyclone is likely
449 to be excited in this region. In winter, the excitation of plasma vortices is associated with the
450 invasion of cosmic rays into the polar atmosphere. Whirlwinds are "explosive", growing
451 rapidly, like tornadoes, but their average life time is about a day. Polar winter cyclones are
452 intensified during phase transitions of moisture. In the polar latitudes of the Russian northern
453 seas, moist air is transferred from the warm Atlantic and in the east from the Pacific Ocean.
454 Seasonal differences were found in the effect of solar activity on individual types of
455 atmospheric circulation (including reliable correlations) [49].

456 Tornadoes are excited and amplified over the ocean in low latitudes. The moisture from the
457 surface of the ocean evaporates under the influence of the solar photon flux, on pressure
458 gradients it is pumped and condensed on aerosol particles in a tornado. In [50], the effect of
459 cosmic rays on tornado amplification was considered.

460 Aerosol plasma vortices interact with Rossby waves and vortices. The excitation of Rossby
461 waves and vortices in hydrodynamic flows is associated with large-scale pressure gradients
462 in inhomogeneous heating of the atmosphere (and the ocean) and with the dependence of
463 the Coriolis force on the latitude. The effect of Coriolis force associated with the Earth's
464 rotation on atmospheric movements is estimated by the ratio of inertial force to Coriolis
465 force. The Arctic region of the atmosphere is characterized by maximum Coriolis force due
466 to the Earth's rotation, and the Rossby number for winter polar cyclones is of order $Ro \leq 10$
467 [18,51].

468 For the vortex structures of the neutral atmosphere, the Obukhov-Charney equation used
469 the approximation of shallow water rotating at a constant speed, with a free surface in the
470 gravitational field and a given bottom relief. In the coordinate system with the z axis directed
471 along the local vertical, x - to the east, y - to the north, the equation describes the
472 preservation of the vortex [51]

$$473 \quad \frac{\partial}{\partial t}(\Delta \Psi - L_0^{-2} \Psi) + \{ \Psi, \Delta \Psi \} + \beta \frac{\partial \Psi}{\partial x} - \gamma \frac{\partial \Psi}{\partial y} = 0, \quad (7)$$

474 here $\{ \dots \}$ - Poisson brackets, $\Delta = \partial^2 / \partial x^2 + \partial^2 / \partial y^2$, $u = -\partial \Psi / \partial y$, $v = \partial \Psi / \partial x$, $\Psi = gh / f_0$,
475 $\beta = (f_0 / H_0)(\partial h_1 / \partial y)$, $\gamma = (f_0 / H_0)(\partial h_1 / \partial x)$, $f = f_0(1 + h_1 / H_0)$, $f_1 = 2\Omega_0 \sin(\phi)$ - is Coriolis
476 parameter, Ω_0 - is the Earth rotation speed, ϕ - is latitude, $f_0 = f_1(45^\circ)$, $L_0 = \sqrt{gH_0} / f_0$. The

477 function $h_1(x, y, t)$ is set to the relief, the unevenness of the bottom. The parameters β, γ
478 are responsible for the dispersion of large-scale wave processes. The effect of a non-ionized
479 aerosol impurity during condensation of moisture on vortices and Rossby waves could be
480 inscribed in the equation like the effect of bottom relief. The gyrotropy of an aerosol ionized
481 impurity in the geomagnetic field, the excitation of plasma vortices increase atmospheric
482 vortex activity. In equation (7) the effect of the MHD generator is not marked. The interaction
483 of Rossby vortices and plasma vortices occurs at the rotor level as the interaction of particle
484 velocity vortices.

485

486 **3. CONCLUSION**

487

488 Under the influence of atmospheric vortex structures, jet streams become sinuous. The jet
489 flow directed northward forms in the region of the approach of the Atlantic cyclone and
490 anticyclone. Anticyclone accumulates pollution. The jet flow transfers pollution and moist
491 cyclonic air into high latitudes, and under its pressure the polar flow acquires a bend. A polar
492 cyclone is likely to be excited in this area.

493 In winter, the excitation of plasma vortices is associated with the invasion of cosmic rays into
494 the polar atmosphere. The ionization of an aerosol impurity by particles of cosmic rays
495 accelerates the processes of condensation and crystallization of atmospheric moisture. An
496 increase in the concentration of aerosols and cosmic ray fluxes of solar and galactic origin
497 causes a nonlinear in power response of atmospheric processes, increasing the excitation of
498 plasma vortices. Nonlinearity is associated with the cascade nature of the ionization process,
499 with the effect of condensation of moisture and latent heat upon excitation and enhancement
500 of aerosol plasma vortices in the atmosphere.

501 At low latitudes, powerful atmospheric vortices, tornadoes, are excited over the ocean. In the
502 high-latitude troposphere in the region of the geomagnetic polar cap in the winter period,
503 local cyclonic structures with ice storms, invasions into middle and even subtropical latitudes
504 are excited. The vortices are explosive in nature, growing rapidly. The time of excitation of
505 such cyclones is 15–25 hours, which is comparable with the time of initiation of a tornado.
506 The nature of the vortex is electromagnetohydrodynamic. The accumulation of energy of
507 cyclonic structures is associated with moisture condensation in mosaic cellular distributions
508 of aerosol impurities. When aerosols are ionized by solar photons and cosmic rays, plasma
509 vortices are generated in the cells in the geomagnetic field. During the interaction of plasma
510 vortices and Rossby vortices, a large-scale vortex structure is formed and grows.

511 The formation of a mosaic cellular distribution of aerosols is associated with the amplification
512 of acoustic-gravity wave packets at wind speed gradients. At low latitudes, excitation of drift
513 instabilities in an aerosol plasma is likely. The formation of the cellular structure is
514 associated with the electrostatic instability of an inhomogeneous plasma, an aperiodic
515 increase of the electric field in the plasma inhomogeneity. In plasma flows orthogonal to the
516 geomagnetic field, an MHD electric field generator is excited, and the electric field in plasma
517 inhomogeneities is amplified. Pollution, aerosol impurities contribute to the excitation and
518 strengthening of atmospheric vortices. The influence of anthropogenic pollution on
519 atmospheric processes is increasing. With the intensification of the vortex structures, the
520 climate contrast, floods, droughts, hurricanes, and the invasion of cold in low latitudes
521 increase.

522 During attenuation of acoustic-gravity waves on the wind speed gradients, a mosaic cellular
523 structure of the distribution of the aerosol impurity is formed. The excitation of plasma
524 vortices in cells and their electromagnetic interaction accelerates the growth of a large-scale
525 vortex. Plasma vortices at the rotor level interact with Rossby vortices. In the genesis of

526 vortices, aperiodic electrostatic disturbances of an inhomogeneous plasma play an important
527 role.

528

529 **COMPETING INTERESTS**

530

531 Authors have declared that no competing interests exist.

532

533 **AUTHORS' CONTRIBUTIONS**

534

535 *This work was carried out in collaboration between all authors. All authors read and*
536 *approved the final manuscript.*

537

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