

3  
4 **Oblique Propagation of Nonlinear Solitary Waves in**  
5 **Magnetized Plasma with Nonextensive Electrons**  
6

7  
8  
9  
10  
11 **ABSTRACT**  
12

In this paper, the properties of obliquely propagating nonlinear solitary waves have been investigated theoretically under the effect of magnetic field. The plasma system is considered to be consisting of nonextensively distributed electrons and stationary ions. The nonlinear Korteweg-de-Vries (KdV) equation and its solution have been derived by using reductive perturbation method. Effect of various parameters such as electron nonextensivity, external magnetic field and obliqueness on the properties of solitary waves is investigated numerically. A critical value of nonextensivity is found at which solitary structures transit from negative to positive potential. The numerical results are interpreted graphically. The results may be useful for understanding the wave propagation in laboratory and space plasmas where magnetic field is present.

13  
14 *Keywords: Magnetized plasma, q-nonextensive distribution, reductive perturbation method,*  
15 *nonlinear waves and soliton*

16  
17 **1. INTRODUCTION**

18 The nonlinear wave structures have provided a fascinating research field for plasma  
19 physics community due to their importance in explaining various laboratory, space and  
20 astrophysical atmosphere [1-3]. Nonlinear structures like solitons, shock waves, double  
21 layers etc. are observed both in space and laboratory. Out of them, solitons have become a  
22 main source of interest for researchers from across the globe owing to their rich physical  
23 insight underlying the various nonlinear phenomena. Solitons are stable nonlinear entities  
24 that arise due to delicate balance of nonlinearity and dispersion. Nonlinear wave structures  
25 in various plasma models and compositions have been investigated for the last half century,  
26 both theoretically and observationally [4-8]. There exists a strong magnetic field on the  
27 surface of fast rotating neutron stars and in the pulsar magnetosphere [9-10] which has a  
28 significant impact on the nonlinear wave propagation. Considering this, an immense interest  
29 has been developed in researchers to study nonlinear propagation of ion-acoustic waves in  
30 magnetized plasmas [11-16]. The nonlinear propagation of the electron-acoustic (EA) waves  
31 in magnetized plasma has been considered by Dubouloz et al [17]. They reported that the  
32 electric field spectrum produced by an Electron-acoustic solitary wave (EASW) is not  
33 significantly modified by the presence of a magnetic field. Mace and Hellberg [18] studied  
34 the influence of the magnetic field on the features of the weakly nonlinear electron-acoustic  
35 waves in magnetized plasma. They have predicted the existence of negative potential  
36 structures in both magnetized and unmagnetized cases. Devanandhan et al [19] have  
37 investigated electron-acoustic solitary waves (EASWs) in two component magnetized  
38 plasma and predicted negative solitary potential structures. They further showed that with  
39 the increase in magnetic field, the soliton electric field amplitude increases while the soliton

40 width and pulse duration decreases. The properties of small amplitude wave in magnetized  
 41 plasma are investigated by Pakzad and Javidan [20]. It was found that both rarefactive  
 42 (negative amplitude) and compressive (positive amplitude) solitons can be propagated and  
 43 soliton profiles became narrower in stronger magnetic fields.

44 The deviations of electron populations from their thermodynamic equilibrium have been  
 45 reported by many space plasma observations. A nonextensive distribution is the most  
 46 generalized distribution to study the linear and nonlinear properties of solitary waves in  
 47 different plasma systems, where the non-equilibrium stationary states exist. The  
 48 nonextensive statistical mechanics has gathered immense attention over the last two  
 49 decades. This mechanics is based on the deviations of Boltzmann-Gibbs-Shannon (B-G-S)  
 50 entropy measures first recognized by Renyi [21] subsequently proposed by Tsallis [22]. The  
 51 Maxwellian distribution in Boltzmann-Gibbs statistics is valid universally for the macroscopic  
 52 ergodic equilibrium systems. While for systems having long-range interactions, the complete  
 53 description of the features becomes inadequate with Maxwellian distribution. The parameter  
 54  $q$  that underpins the generalized entropy of Tsallis. Further,  $q$  is associated to the underlying  
 55 dynamics of the system and measures the amount of its nonextensivity. Generalized entropy  
 56 of whole is greater (smaller) than the entropies of subsequent parts if  $q < 1$  i.e.  
 57 superextensivity ( $q > 1$  i.e. subextensivity). The nonextensive statistics has found applications  
 58 in a large quantity of astrophysical and cosmological atmospheres such as stellar polytropes  
 59 [23], the solar neutrino problem [24], peculiar velocity distributions of galaxies [25] and  
 60 systems with long range interactions and also fractal-like space-times. Different types of  
 61 waves, viz. ion acoustic (IA) waves, electron-acoustic (EA) waves, or dust-acoustic (DA)  
 62 waves, in nonextensive plasmas are investigated by many researchers considering one or  
 63 two components to be nonextensive [26-34]. In the present investigation, we aim at studying  
 64 the obliquely propagating solitary waves in magnetized plasma system with nonextensive  
 65 distributed electrons. The paper is organized as follows: in Sec. 2, the basic equations  
 66 governing the plasma dynamics and their analysis are given. In Sec. 3, we present the  
 67 numerical analysis and discussion of the results. Finally, we conclude the paper in Sec. 4.

68  
 69

## 2. BASIC EQUATIONS AND NONLINEAR ANALYSIS

70 Let us consider the homogeneous magnetized plasma containing  $q$ -nonextensive electrons  
 71 and stationary ions. The external static magnetic field is assumed to point in the  $z$ -direction  
 72 i.e.  $B = B_0 \hat{z}$ . The dynamics of the propagation of waves in such magnetized plasma is  
 73 governed by the following set of normalized equations:

$$74 \quad \frac{\partial n}{\partial t} + \nabla \cdot (nu) = 0 \quad (1)$$

$$75 \quad \frac{\partial u}{\partial t} = (u \cdot \nabla)u = -\nabla \phi - \omega_0 (u \times \hat{z}) - \frac{5}{3} \frac{\sigma}{n^{1/3}} \nabla n \quad (2)$$

$$76 \quad \nabla^2 \phi = n_e - n \quad (3)$$

77

78 where  $n$  and  $u$  are the ion number density and ion fluid velocity normalized to equilibrium  
 79 plasma density  $n_0$  and ion acoustic speed  $C_s = (T_e/m)^{1/2}$ ,  $T_e$  is the electron temperature and  $m$   
 80 is the mass of positively charged ions, respectively.  $\phi$  is the electrostatic wave potential  
 81 normalized to  $T_e/e$ , where  $e$  is the magnitude of electron charge and  $\sigma = T_i/T_e$  with  $T_i$  being the  
 82 ion temperature. In this plasma model, ion plasma period  $\omega_p^{-1} = (m/4\pi n_0 e^2)^{1/2}$ , the Debye  
 83 length  $\lambda_D = (T_e/4\pi n_0 e^2)^{1/2}$  and ion cyclotron frequency is given by  $\omega_c = (eB_0/m\omega_p)$ . The number  
 84 density of electron fluid with nonextensive distribution is given by:

85 
$$n_e = (1 + (q-1)\phi)^{\frac{(q+1)}{2(q-1)}} \quad (4)$$

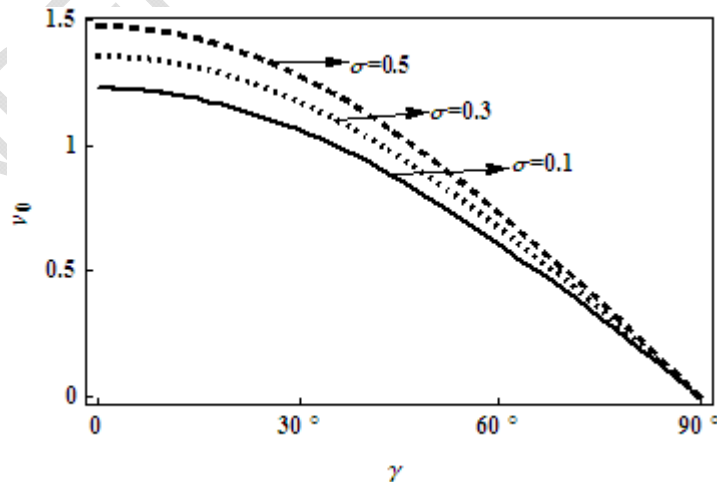
86  
87 where  $q$  is the nonextensivity parameter. The electron distribution reduces to the well-known  
88 Maxwell Boltzmann distribution for the extensive limiting case  $q$  approaches to 1 [34]. In  
89 transformations given by Gardner and Morikawa [35] put  $\alpha=1/2$  the stretched coordinates  
90 becomes  $\xi=\varepsilon^{1/2}(l_x x+l_y y+l_z z-v_0 t)$ ,  $\tau=\varepsilon^{3/2}t$ . Here  $v_0$  is the linear phase velocity and  $\varepsilon$  is a small  
91 parameter.  $l_x, l_y, l_z$  are the direction cosines of the wave vector with respect to the  $x, y$  and  $z$   
92 axes respectively. The perturbed quantities are expanded in power series of  $\varepsilon$  as follows:  
93

94 
$$\begin{aligned} n &= 1 + \varepsilon n^{(1)} + \varepsilon^2 n^{(2)} + \varepsilon^3 n^{(3)} + \dots \\ u_{x,y} &= 0 + \varepsilon^{\frac{3}{2}} u_{x,y}^{(1)} + \varepsilon^2 u_{x,y}^{(2)} + \varepsilon^{\frac{5}{2}} u_{x,y}^{(3)} + \dots \\ u_z &= 0 + \varepsilon u_z^{(1)} + \varepsilon^2 u_z^{(2)} + \varepsilon^3 u_z^{(3)} + \dots \\ \phi &= 0 + \varepsilon \phi^{(1)} + \varepsilon^2 \phi^{(2)} + \varepsilon^3 \phi^{(3)} + \dots \end{aligned} \quad (5)$$

95  
96 Now using the number density of electron fluid given by equation (4), stretching coordinates  
97  $\xi$  and  $\tau$  and the expansions (5) into (1)-(3). Comparing the coefficients of lowest order of  $\varepsilon$   
98 i.e.  $\varepsilon^{3/2}$ , we get the linear dispersion relation which is given by the following expression.  
99

100 
$$v_0^2 = \frac{l_z^2}{c_1} \left[ 1 + \frac{5}{3} \sigma c_1 \right] \quad (6)$$

101  
102 where  $c_1=(q+1)/2$  and the phase velocity depends upon the ion to electron temperature ratio  
103  $\sigma$ , the strength of nonextensivity  $q$  and obliqueness of propagation  $\gamma$ . Mathematical relation  
104 (6) shows that phase velocity increases with ion to electron temperature ratio  $\sigma$  and  
105 decreases with non-extensive parameter ( $q$ ) for all ranges of  $q$ . The  $q$ -dependence of phase  
106 velocity occurs from the factor  $c_1$  in the expression (6). Similar kind of behavior has been  
107 observed by Sahoo et al [36] and Akhtar et al [28] in their respective researches.



108  
109 **Fig.1. Variation of wave phase velocity ( $v_0$ ) with angle of propagation ( $\gamma$ ) for  $\sigma= 0.10$**   
110 **(Solid Line), 0.30 (Dotted Line), 0.50 (Dashed Line) with  $q= 0.5$ .**

111  
 112 Figure 1 shows the typical variation of the phase velocity  $v_0$  with respect to angle of  
 113 propagation  $\gamma$  for three different values of ion to electron temperature ratio  $\sigma$ . It is observed  
 114 that wave phase velocity decreases with angle between the direction of the wave  
 115 propagation vector  $k$  and the external magnetic field  $B_0$ . The decrease of  $v_0$  with  $\gamma$  also  
 116 becomes clear from the expression (6) where  $v_0 \propto \sqrt{\cos \gamma}$  and becomes zero for  $\gamma = 90^\circ$ .  
 117 This decreasing trend of  $v_0$  with  $\gamma$  is similar to that observed by Misra and Wang [37]. It is  
 118 also clear from the figure that phase velocity increases with increase in temperature ratio  $\sigma$ .  
 119 Further, the wave phase velocity is found to be independent of the magnetic field strength  
 120 and decreases with nonextensivity  $q$ .

121  
 122 Going to the next higher order of  $\varepsilon$  i.e.  $\varepsilon^2$  and by doing algebraic manipulations, we get the  
 123 following Korteweg-de Vries (KdV) equation (7) in which we have replaced  $\phi^{(1)}$  with  $\phi$   
 124 for simplicity.

$$125 \quad \frac{\partial \phi}{\partial \tau} + A \phi \frac{\partial \phi}{\partial \xi} + B \frac{\partial^3 \phi}{\partial \xi^3} = 0 \quad (7)$$

126 where A is non-linear and B is dispersion coefficients and are given as:

$$127 \quad A = l_z \sqrt{c_1} \sqrt{1 + \frac{5}{3} \sigma c_1} \left[ \frac{3}{2} - \left[ \frac{5\sigma}{18} + \frac{c_2}{c_1^3} \right] \left[ \frac{c_1}{1 + \frac{5}{3} \sigma c_1} \right] \right] \quad (8)$$

$$128 \quad B = \frac{1}{2} \frac{l_z}{c_1 \sqrt{c_1} \sqrt{1 + \frac{5}{3} \sigma c_1}} \left[ 1 + \left[ \frac{1 - l_z^2}{\omega_0^2} \right] \left[ 1 + \frac{5}{3} \sigma c_1 \right]^2 \right] \quad (9)$$

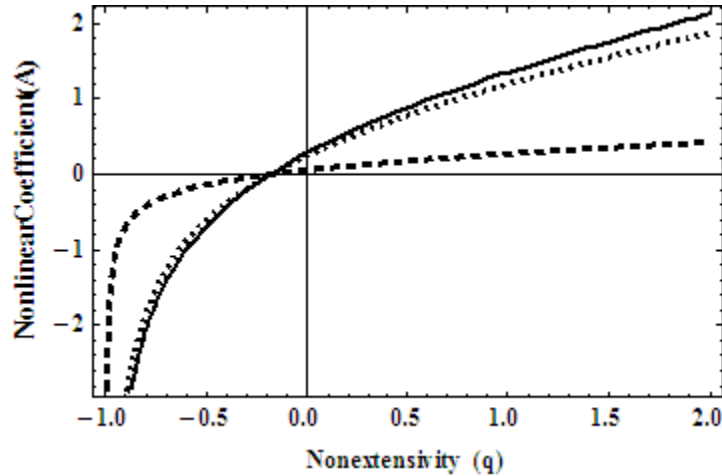
129  
 130 The stationary solitary wave solution of Eq. (7) is directly given by  
 131

$$132 \quad \phi = \phi_0 \left[ \sec h \left( \frac{\eta}{\delta} \right) \right]^2 \quad (10)$$

133  
 134 where the amplitude  $\phi_0$  and width  $\delta$  of the soliton are given by  $\phi_0 = 3u_0/A$  and  $\delta = (4B/u_0)^{1/2}$   
 135 and here  $\eta = \xi - u_0 \tau$ . From the expressions of A and B (Eqns. (8) and (9)), it is found that  
 136 the amplitude of the soliton does not depend on the external magnetic field but depends on  
 137 the ion and electron temperature ratio  $\sigma$ . On the other hand, the width of the soliton depends  
 138 on the strength of external magnetic field.

### 139 3. RESULTS AND DISCUSSION

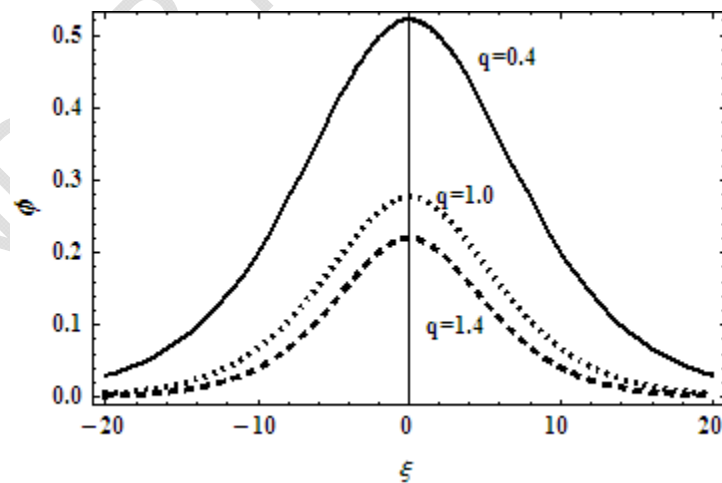
141 We have investigated the effects of  $q$ -nonextensive electrons, magnetic field and angle  
 142 of propagation on the nonlinear wave propagation of solitary waves in magnetized plasma.  
 143 To describe the nonlinear propagation of the waves, we have derived a KdV equation (7)  
 144 and obtained solitary wave solution (10). Depending upon the value of nonlinear coefficient  
 145 A, the solitary wave might be associated with positive or negative potentials. Equation (8)  
 146 indicates that A is dependent on parameters  $q, \sigma, l_z = \cos(\gamma)$  which define the nature of  
 147 solitary waves.



148  
149

150 **Fig.2. Variation of nonlinear coefficient (A) with nonextensivity q for  $\gamma = 0^\circ$  (Solid Line),**  
151  **$40^\circ$  (Dotted Line),  $80^\circ$  (Dashed Line) with  $\sigma = 0.50$ .**

152 The nonlinear coefficient (A) as a function of nonextensivity (q) is displayed in Figure 2 for  
153 three different values of angle  $\gamma$ . A transition from negative to positive potential structures  
154 results with increase in non-extensive parameter (q). A negative critical value  $q_c$  is obtained  
155 for a fixed set of parametric values. We observe that at  $q > q_c$ , positive (hump shape or  
156 commonly known as compressive soliton) solitary waves exist, whereas at  $q < q_c$ , negative  
157 (dip shape or rarefactive solitons) solitary waves exist. It may be further mentioned that the  
158 critical value  $q_c$  remains same for all values of  $\gamma$ . From Eq. (8), it can be seen that  
159 magnitude of external magnetic field ( $B_0$ ) has no influence on the nonlinear coefficient A. In  
160 order to investigate the effect of nonextensive parameter q, solitary wave profiles for  $0 < q < 2$   
161 and  $-1 < q < 0$  have been displayed in Figures 3 and 4 respectively. As  $q = 1$  corresponds to  
162 Maxwellian behavior, hence increase in nonextensivity leads to change in the amplitude of

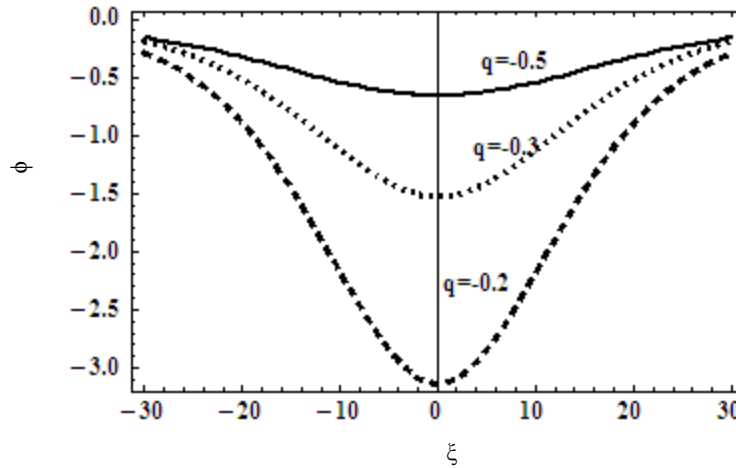


163  
164  
165  
166  
167  
168

**Fig. 3. Variation of soliton solution  $\phi$  as a function of  $\xi$  for three different values of q=**  
**0.4 (Solid Line), 1.0 (Dotted Line) and 1.4 (Dashed Line) with  $\gamma = 30^\circ$ ,  $\sigma = 0.20$ ,  $\omega_0 = 0.40$**   
**and  $u_0 = 0.10$ .**

169 the solitary structures. It can be interpreted that towards the higher value of nonextensivity  
 170 from the Maxwellian i.e.  $q=1$ , peak amplitude decreases while it increases while it increases

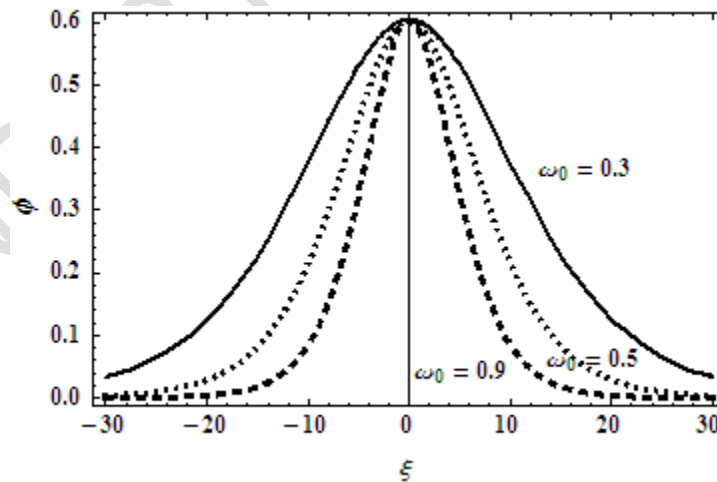
171



172

173 **Fig.4. Variation of soliton solution  $\phi$  as a function of  $\xi$  for different values of  $q = - 0.5$**   
 174 **(Solid Line),  $- 0.3$  (Dotted Line) and  $- 0.2$  (Dashed Line) with  $\gamma = 60^\circ$ ,  $\sigma = 0.20$ ,  $\omega_0 = 0.40$**   
 175 **and  $u_0 = 0.10$ .**

176 towards lower nonextensivity side ( $q=0.5$ ). In other words, the amplitude of the positive  
 177 (negative) KdV soliton is seen to decrease with the increasing values of nonextensivity. This  
 178 behavior of soliton amplitude is similar to that observed by Pakzad and Javidan [20], Akhtar  
 179 et al [28] and Sahoo et al [36]. It is observed that width of both the solitons decreases with  
 180 increase in  $q$ , which are in agreement with the results of Pakzad and Javidan [20] and Akhtar  
 181 et al [28] but contrary to Sahoo et al [36].



182

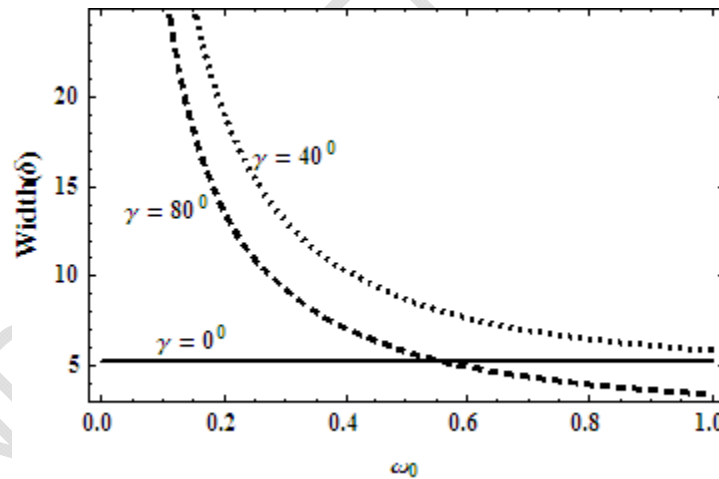
183

184

185 **Fig.5. Variation of soliton solution  $\phi$  as a function of  $\xi$  for different values of  $\omega_0 = 0.30$**   
 186 **(Solid Line),  $0.50$  (Dotted Line) and  $0.90$  (Dashed Line) with  $q = 0.5$ ,  $\gamma = 50^\circ$ ,  $\sigma = 0.20$  and**  
 187  **$u_0 = 0.10$ .**

188 The ion cyclotron frequency ( $\omega_0$ ) characterizes the influence of the magnetic field strength on  
 189 the properties of solitons. Figure 5 shows the variation of the soliton solution  $\phi$  with  
 190 parameter  $\xi$  for three different values of ion cyclotron frequency  $\omega_0$ . It has been seen that  
 191 the amplitude remains same while the width of the soliton decreases on applying stronger  
 192 magnetic field. It is due to the fact that the radius of particle circular motion is reduced due to  
 193 the greater magnetic field. From expression of amplitude, it is also evident that amplitude of  
 194 the solitary waves does not influenced by the intensity of the external magnetic field  $B_0$ . The  
 195 width of the soliton is influenced by the ion cyclotron frequency through the coefficient B in  
 196 the expression of width. Thus the effect of external magnetic field is to make the soliton  
 197 spiky, which is in agreement with the results obtained by Sahoo et al [36] and Misra and  
 198 Wang [37]. A similar kind of behavior is observed for the range of nonextensivity  $-1 < q < 0$  (not  
 199 shown).

200 Further, the angle of propagation ( $\gamma$ ) in the presence of magnetic field also plays an  
 201 important role in soliton mechanism. In order to understand this behavior, a plot of soliton  
 202 width  $\delta$  as a function of  $\omega_0$  has been given in figure 6 taking three different values of  $\gamma$  and  
 203 other parameters are given in caption. From the figure 6, it becomes clear that  $\gamma=0^\circ$  has no  
 204 effect on the width of soliton as indicated by a straight line (solid line). However, for  $\gamma \neq 0^\circ$ , the  
 205 behavior changes and magnetic field start playing its role. Dotted ( $\gamma=40^\circ$ ) and dashed lines  
 206 ( $\gamma=80^\circ$ ) show that width starts decreasing with the application of magnetic field. This  
 207 behavior sets in early for larger  $\gamma$ . Similar result was predicted by Pakzad and Javidan [20].  
 208 It is further observed that ion to electron temperature ratio ( $\sigma$ ) has stronger influence on the  
 209 soliton structures. It is evident that higher magnitudes of  $\sigma$  cause significant reduction in the  
 210 amplitude of the solitary waves. But the soliton width increases with increase in  $\sigma$ , a result

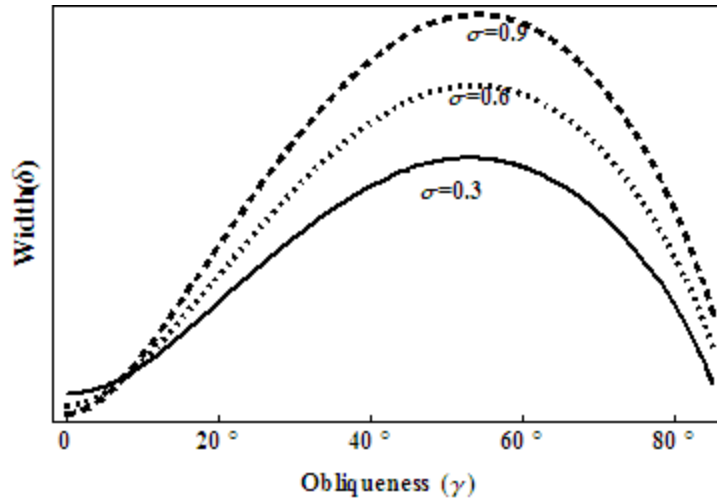


211  
 212  
 213  
 214  
 215

**Fig.6. Width of soliton ( $\delta$ ) as a function of  $\omega_0$  for different values of  $\gamma = 0^\circ$  (Solid Line),  $40^\circ$  (Dotted Line) and  $80^\circ$  (Dashed Line) with  $\sigma = 0.20$ ,  $\omega_0 = 0.40$  and  $u_0 = 0.10$ .**

216  
 217  
 218  
 219  
 220  
 221  
 222  
 223

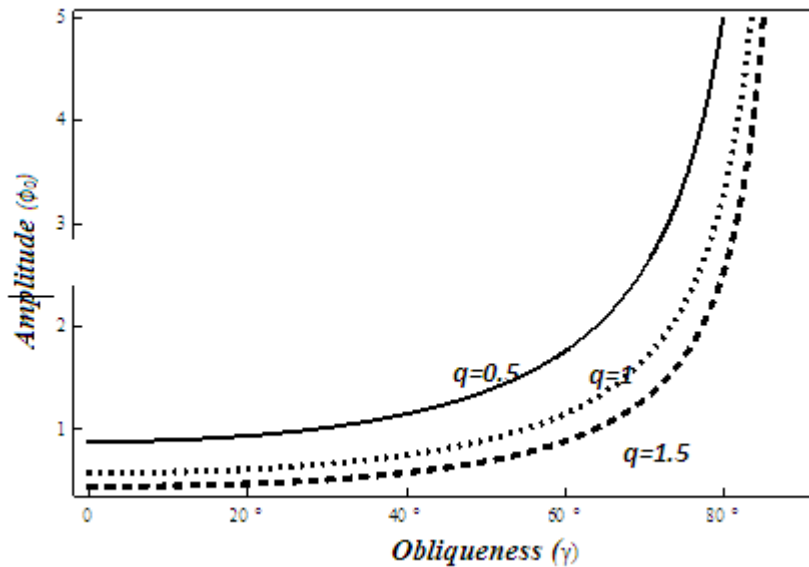
which is in agreement with Akhtar et al [28]. Figure 7 presents soliton width  $\delta$  as a function of angle of propagation ( $\gamma$ ) with respect to magnetic field with three different values of ion to electron temperature ratio  $\sigma$ . This figure clearly shows that for the lower limit ( $0^\circ$ - $50^\circ$ ) of angle of propagation, the width of the positive (negative) KdV soliton increases. But for the higher limits ( $50^\circ$ - $90^\circ$ ) of angle of propagation, width of both the solitons decreases. Also, it is clear from this figure that as  $\gamma$  approaches to  $90^\circ$ , the width of the solitary waves approaches to zero. Similar results were obtained by Haider et al [38] and Pakzad and Javidan [20].



224  
225  
226  
227

**Fig.7. Variation of soliton width  $\delta$  as a function of  $\gamma$  for different values of  $\sigma = 0.30$  (Solid Line),  $0.60$  (Dotted Line) and  $0.90$  (Dashed Line) with  $q = 0.5$ ,  $\omega_0 = 0.40$  and  $u_0 = 0.10$ .**

228 To study the effects of the angle of propagation of the wave amplitude of soliton wave profile  
229 has been plotted as a function of  $\gamma$  in Figure 8. For the positive nonextensivity  $0 < q < 2$ , the  
230 amplitude of the compressive soliton increases with the increase in obliqueness of the wave  
231 propagation  $\gamma$ . An opposite trend has been observed for the negative nonextensivity i.e  
232 amplitude decreases with the increase in  $\gamma$  (not shown here). It may be mentioned that our  
233 results are similar to Pakzad and Javidan [20] for the range  $-1 < q < 0$ , however, for the range  
234  $0 < q < 2$ , these results show an opposite trend. Pakzad and Javidan [20] showed that  
235 amplitude of the solitary waves decreases with increase in angle of propagation for both the  
236 ranges  $-1 < q < 0$ , and  $0 < q < 2$ .



237  
238  
239  
240  
241

**Fig.8. Variation of soliton amplitude  $\phi_0$  as a function of  $\gamma$  for different values of  $q = 0.5$  (Solid Line),  $1.0$  (Dotted Line) and  $1.5$  (Dashed Line) with  $\sigma = 0.50$ ,  $\omega_0 = 0.40$  and  $u_0 = 0.30$ .**



242 . **4. CONCLUSION**

243 The q-nonextensive electrons, strength of magnetic field, ion to electron temperature  
244 ratio and obliqueness of wave propagation significantly change the solitary structures. In nut  
245 shell, our main findings are summarized below:

246 (i) Phase velocity of solitary wave decreases with increase in the q-nonextensive  
247 parameter and angle of propagation but increases with ion to electron temperature ratio.

248 (ii) There exists a critical value  $q_c$  for a fixed set of parametric values. For  $q > q_c$ , positive  
249 (hump shape) solitary waves exist, whereas at  $q < q_c$ , negative (dip shape) solitary waves  
250 exist.

251  
252 (iii) The amplitude of the positive (negative) KdV soliton is seen to decrease with the  
253 increasing values of nonextensivity. It is also observed that width of both kind of solitons  
254 decreases with increase in  $q$ .

255 (iv) The amplitude of solitary structures is independent of the applied magnetic field however  
256 the width decreases with magnetic field.

257  
258 (v) For the lower limit ( $0^\circ - 50^\circ$ ) of angle of propagation, the width of the positive (negative)  
259 KdV soliton increases. But for the higher limits ( $50^\circ - 90^\circ$ ) of angle of propagation, width  
260 decreases.

261 (vi) The amplitude of the positive KdV soliton increases with the increase in obliqueness of  
262 the wave propagation  $\gamma$ . On the other hand, the amplitude of the negative KdV soliton  
263 decreases with the increase in  $\gamma$ .

264  
265 The present investigation may help us to understanding the study of nonlinear waves in  
266 astrophysical plasmas. The comparison of the obtained results with other particle  
267 distributions can help us to find better knowledge in plasma physics.

268  
269 **REFERENCES**

270 1. Ikezawa S, Nakamura Y. Observation of Electron Plasma Waves in Plasma of Two-  
271 Temperature Electrons. J Phys Soc Jpn. 1981;50: 962-67.

272 2. Dubouloz N, Pottellette R, Malingre M, Holmgren G, Lindqvist P A. Detailed analysis of  
273 broadband electrostatic noise in the dayside auroral zone. J Geophys Res. 1991;96:3565.

274 3. Mozer F S, Ergun R, Temarin M, Cattell C, Dombeck J, Wygant J. New Features of Time  
275 Domain Electric-Field Structures in the Auroral Acceleration Region. Phys Rev  
276 Lett.1997;79:1281.

277 4. Stasiewicz K. Nonlinear Alfvén, magnetosonic, sound, and electron inertial waves in fluid  
278 formalism. J Geophys Res. 2005;110:A03220.

279 5. Shinsuke I, Yukiharu O. Nonlinear Waves along the Magnetic Field in a Multi-Ion Species  
280 Plasma. J Plasma Fusion Res. 2001;4: 500-504.

281 6. Dubinin E M, Sauer K, McKenzie J F, Chanteur G. Nonlinear waves and solitons  
282 propagating perpendicular to the magnetic field in bi-ion plasma with finite plasma pressure.  
283 Nonlinear Process Geophys.2002;9:87-99.

- 284 7. El-Taibany W F, Moslem W M. Higher-order nonlinearity of electron-acoustic solitary  
285 waves with vortex-like electron distribution and electron beam. *Phys Plasmas*.  
286 2005;12:032307.
- 287 8. Gill T S, Bala P, Kaur H, Saini N S, Bansal S. Ion Acoustic Solitons and Double Layers in  
288 a multicomponent plasma consisting of positive and negative ions with nonthermal electrons.  
289 *Eur Phys J D*. 2004;31:91.
- 290 9. Miller H R, Witta P J. *Active Galactic Nuclei*. Springer, Berlin, Germany; 1978.
- 291 10. Michel F C. Theory of pulsar magnetospheres. *Rev Mod Phys*.1982;54:1-66.
- 292 11. Singh S V, Devanandhan S, Lakhina G S, Bharuthram R. Effect of ion temperature on  
293 ion-acoustic solitary waves in a magnetized plasma in presence of superthermal electrons.  
294 *Phys Plasmas*. 2013;20: 012306.
- 295 12. Alinejad H, Mamun A A. Oblique propagation of electrostatic waves in a magnetized  
296 electron-positron-ion plasma with superthermal electrons. *Phys Plasmas*. 2011;18:112103.
- 297 13. Fedousi M, Sultana S, Mamun A A. Oblique propagation of ion-acoustic solitary waves in  
298 a magnetized electron-positron-ion plasma. *Phys Plasmas*. 2015;22:032117.
- 299 14. Mahmood S, Mushtaq A, Saleem H. Ion acoustic solitary wave in homogeneous  
300 magnetized electron-positron-ion plasmas. *New J Phys*. 2003;5:28.1–28.10.
- 301 15. Jehan N, Salahuddin M, Saleem H, Mirza A M. Modulation instability of low-frequency  
302 electrostatic ion waves in magnetized electron-positron-ion plasma. *Phys Plasmas*.  
303 2008;15:092301.
- 304 16. Mio J, Ogino T, Minami K, Takeda S. Modulational instability and envelope solitons for  
305 non-linear Alfvén waves propagating along the magnetic field in plasmas. *J Phys Soc Jpn*.  
306 1976;41:667–73.
- 307 17. Dubouloz N, Treumann R A, Pottelette R, Malingre M. Turbulence generated by a gas of  
308 electron acoustic solitons. *J Geophys Res*. 1993;98:17415–22.
- 309 18. Mace R L, Hellberg M A. The Korteweg–de Vries– Zakharov–Kuznetsov equation for  
310 electron-acoustic waves. *Phys Plasmas*. 2001;8:2649.
- 311 19. Devanandhan S, Singh S V, Lakhina G S, Bharuthram R. Electron acoustic waves in a  
312 magnetized plasma with kappa distributed ions. *Phys Plasmas*. 2012;19:082314.
- 313 20. Pakzad H R, Javidan K. Obliquely propagating electron acoustic solitons in magnetized  
314 plasmas with nonextensive electrons. *Nonlin Process Geophys*. 2013;20:249-55.
- 315 21. Renyi A. On a new axiomatic theory of probability. *Acta Math Hung*. 1955;16:285-335.
- 316 22. Tsallis C. Possible generalization of Boltzmann-Gibbs statistics. *J Stat Phys*.  
317 1988;52:479-87.
- 318 23. Plastino A R. Stellar polytropes and Tsallis' entropy. *Phys Lett A*. 1993;174:384-86.

- 319 24. Kaniadakis G, Lavagno A, Quarati P. Generalized Statistics and Solar Neutrinos. *Phys*  
320 *Lett B*. 1996;369:308-12.
- 321 25. Lavagno A, Kaniadakis G, Rego-Monteiro M, Quarati P, Tsallis C. Non-extensive  
322 thermostistical approach of the peculiar velocity function of galaxy clusters. *Astrophys Lett*  
323 *Commun*. 1998;35: 449–55.
- 324 26. Rossignoli R, Canosa N. Non additive entropies and quantum statistics. *Phys Lett A*.  
325 1999;281:148-53.
- 326 27. Abe S, Martinez S, Pennini F, Plastino A. Nonextensive thermodynamics relations. *Phys*  
327 *Lett A*.2001;281:126-30.
- 328 28. Akhtar N, El Taibany W F, Mahmood S. Electrostatic double layers in arm negative ion  
329 plasma with nonextensive electrons. *Phys Lett A*. 2013;377:1282-89.
- 330 29. Gill T S, Bala P, Kaur H. Electrostatic wave structures and their stability analysis in  
331 nonextensive magnetized electron-positron-ion plasma. *Astrophys Space Sci*. 2015;357:63.
- 332 30. Reynolds A M, Veneziani M. Rotational dynamics of turbulence and Tsallis. *Phys Lett A*.  
333 2004;327:9-14.
- 334 31. Sattin F. Non-Extensive Entropy from Incomplete Knowledge of Shannon Entropy. *Phys*  
335 *Scr*. 2005;71:443-46.
- 336 32. Wada T. On the thermodynamic stability of Tsallis entropy. *Phys Lett A*. 2002;297:334-  
337 37.
- 338 33. Wu J, Che H. Fluctuation in nonextensive reaction–diffusion systems. *Phys Scr*.  
339 2007;75:722-25.
- 340 34. Tribeche M, Djebarni L, Amour R. Ion acoustic solitary waves in a plasma with a q-  
341 nonextensive electron velocity distribution. *Phys Plasmas*. 2010;17:04211.
- 342 35. Gardner C S, Morikawa G K. Similarity in the asymptotic behaviour of collision free  
343 hydromagnetic waves and water waves. *Courant Institute of Mathematical Sciences Rep*.  
344 *NYO*.1960;9082:1-30.
- 345 36. Sahoo H, Chandra S, Ghosh B. Dust Acoustic Solitary Waves in Magnetized Dusty  
346 Plasma with Trapped Ions and q-Non-extensive Electrons. *Afr Rev Phys*. 2015;10:235-41.
- 347 37. Misra A P, Wang Y. Dust-acoustic solitary waves in magnetized dusty plasma with  
348 nonthermal electrons and trapped ions. *Afr Rev Phys*. 2014;10:0032.
- 349 38. Haider M M, Ferdous T, Duha S S, Mamun A A. Dust-ion-acoustic Solitary Waves in  
350 Multi-component Magnetized Plasmas. *Open Journal of Modern Physics*.2014;1:13-24.
- 351  
352  
353  
354